General Relativity

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1. Free movement on geodesics

- (a) Use the action principle $\delta \int \mathcal{L} dt = 0$ and conservation of energy to show that the movement of a free particle is always given by $\delta \int ds = 0$ (*i.e.* free particles move on a path of extremal length, the *geodesic*).
- (b) Write the equations of motion in last week's question 4 for a free particle on a sphere of fixed radius R.
- (c) Using the conserved conjugate momentum to ϕ , show that the remaining equation of motion can be written as

$$\ddot{\theta} - B^2 \frac{\cos\theta}{\sin^3\theta} = 0, \qquad (1)$$

where $B = const = \sin^2 \theta \dot{\phi}$. In the following, we will try to determine which geometrical form the motion will take.

(d) Show that energy conservation implies

$$\frac{\dot{\theta}^2}{\sin^4\theta\,\dot{\phi}^2} = \frac{A}{B^2} - \frac{1}{\sin^2\theta}\,; \qquad A = const = \frac{2E}{mR^2}\,. \tag{2}$$

(e) Calculate $\frac{d}{d\phi} \cot \theta$ and substitute $u \equiv \cot \theta$ to rewrite (2) as

$$\left(\frac{du}{d\phi}\right)^2 = \frac{A}{B^2} - 1 - u^2.$$
(3)

(f) Show that the solution of (3) can be rewritten in cartesian coordinates as

$$z = \alpha x + \beta y, \qquad x^2 + y^2 + z^2 = R^2,$$
 (4)

where α and β are suitably chosen constants.

What form do the trajectories of a free particle on a sphere take?

2. Riemann Tensor

The Christoffel symbols are *not* tensors, and thus are not suitable to describe a curved geometry in a coordinate-invariant way. The only tensor that can be constructed from the metric and its first and second derivatives is the *Riemann tensor*

$$R^{\lambda}{}_{\mu\nu\kappa} \equiv \frac{\partial\Gamma^{\lambda}{}_{\mu\nu}}{\partial x^{\kappa}} - \frac{\partial\Gamma^{\lambda}{}_{\mu\kappa}}{\partial x^{\nu}} + \Gamma^{\eta}{}_{\mu\nu}\Gamma^{\lambda}{}_{\kappa\eta} - \Gamma^{\eta}{}_{\mu\kappa}\Gamma^{\lambda}{}_{\nu\eta} \,. \tag{5}$$

Through self-contractions we get the *Ricci tensor* $R_{\mu\kappa} \equiv R^{\lambda}{}_{\mu\lambda\kappa}$ and the *curvature* scalar $R \equiv g^{\mu\kappa}R_{\mu\kappa}$.

(a) Using the metric, the Riemann tensor can be made fully covariant:

$$R_{\lambda\mu\nu\kappa} = g_{\lambda\eta}R^{\eta}{}_{\mu\nu\kappa}$$

$$= \frac{1}{2} \left(\frac{\partial^2 g_{\lambda\nu}}{\partial x^{\kappa} \partial x^{\mu}} - \frac{\partial^2 g_{\mu\nu}}{\partial x^{\kappa} \partial x^{\lambda}} - \frac{\partial^2 g_{\lambda\kappa}}{\partial x^{\nu} \partial x^{\mu}} + \frac{\partial^2 g_{\mu\kappa}}{\partial x^{\nu} \partial x^{\lambda}} \right)$$

$$+ g_{\eta\sigma} \left(\Gamma^{\eta}{}_{\nu\lambda} \Gamma^{\sigma}{}_{\mu\kappa} - \Gamma^{\eta}{}_{\kappa\lambda} \Gamma^{\sigma}{}_{\mu\nu} \right) .$$
(6)

Check the symmetry properties $R_{\lambda\mu\nu\kappa} = -R_{\lambda\mu\kappa\nu}$ and $R_{\lambda\mu\nu\kappa} = +R_{\nu\kappa\lambda\mu}$.

- (b) Calculate the components of R^{ℓ}_{mnk} , R_{mk} and the curvature scalar R for a space with coordinates (θ, ϕ) and metric $g_{mn} = diag(a^2, a^2 \sin^2 \theta)$. (Use last week's Christoffel symbols again!)
- (c) Do the same for a space with coordinates (x^1, x^2) and metric $g_{mn} = diag(1, (x^1)^2)$. What is the geometry?
- (d) What is the curvature of a circle with fixed radius R?