

Summary and Outlook

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- 2 Challenges to the SM

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- Nevertheless, model builders are running wild: “hidden valleys”, “unparticles”, “ 10^{32} hidden sectors”, ...
- Phenomenologists are waiting for LHC (used to be waiting for SSC...): work on (N)NLO corrections; improved mass reconstruction techniques; Will be useful *if* LHC finds right kinds of particles.

2 Challenges to the SM

2a: $g_\mu - 2$

- Still differs from SM:

$$a_\mu^{\text{exp}} - a_\mu^{\text{SM, ee data}} = (292 \pm 63 \pm 58) \cdot 10^{-11} \quad \text{PDG}$$

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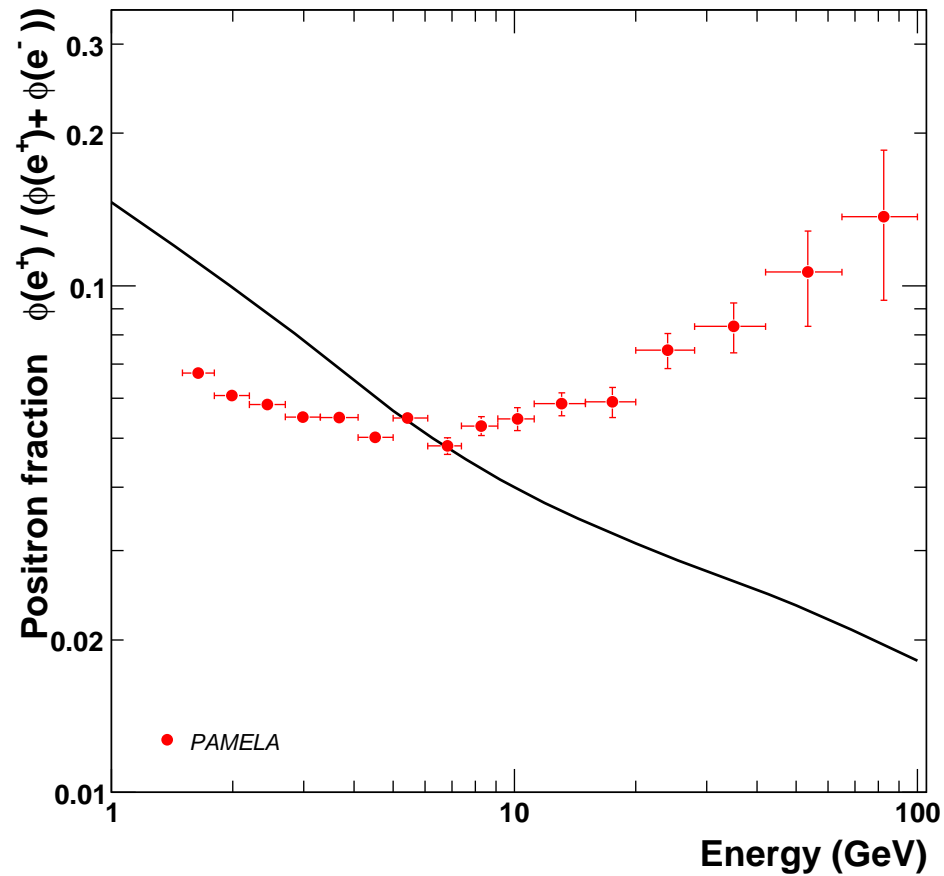
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- New FNAL proposal for next generation expt. (R.M. Carey et al.)

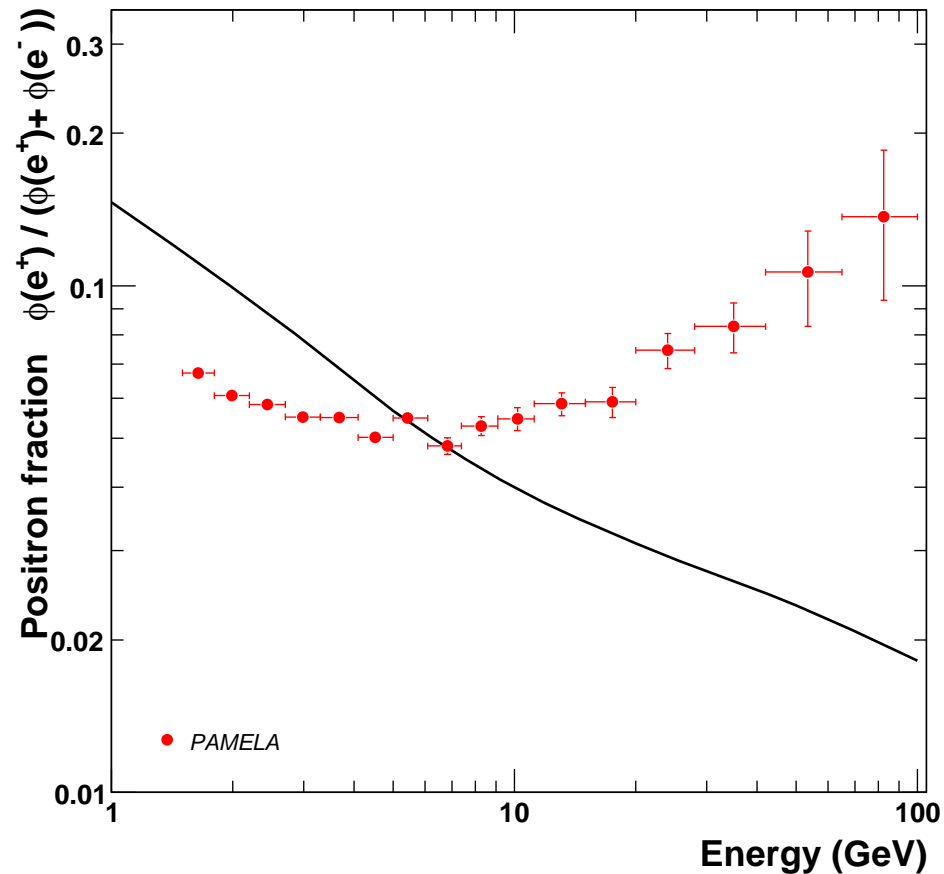
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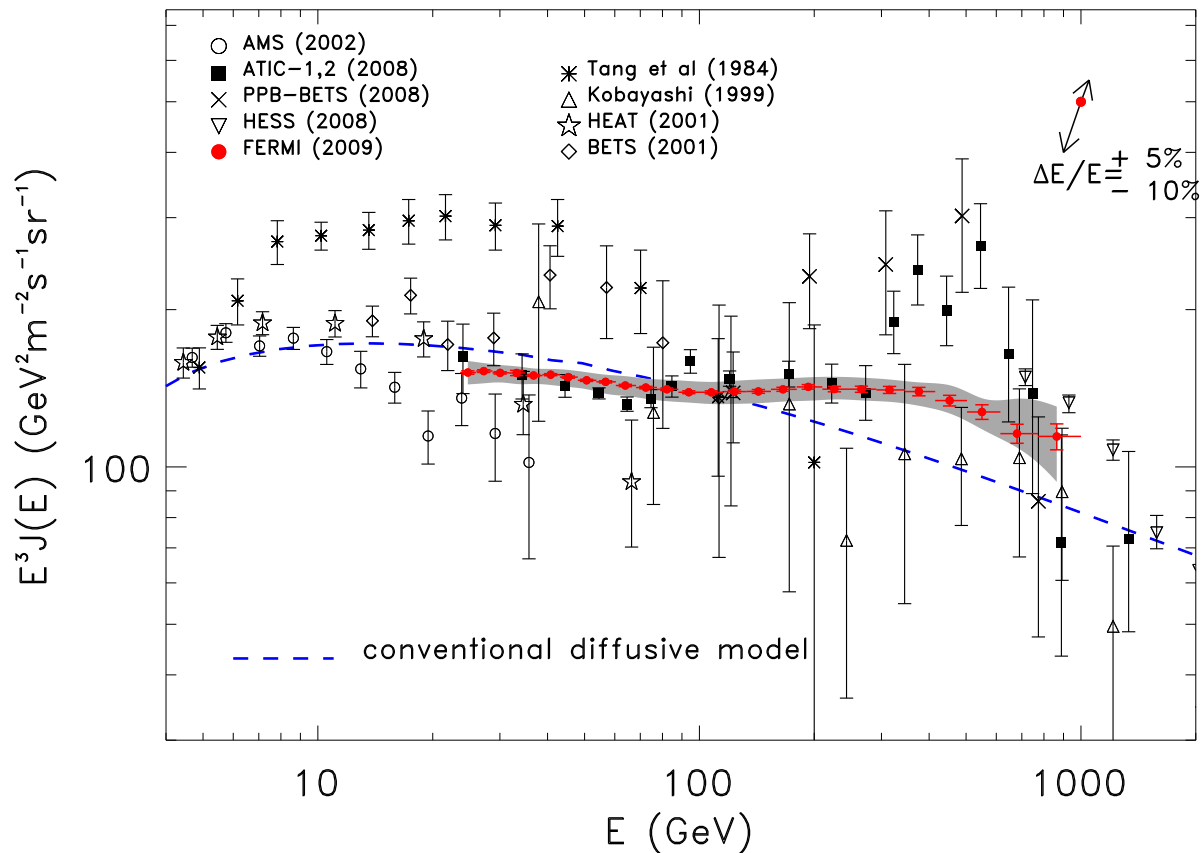
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Very poor agreement between data and “theory” over almost entire energy range!

PAMELA etc. (cont'd)

ATIC, PPB-BETS: Claimed “peak” in $e^+ + e^-$ spectrum
(after multiplying with E^3 !): **Refuted by Fermi-LAT!**



H.E.S.S. finds steeply falling spectrum at $E \gtrsim 1 \text{ TeV}$

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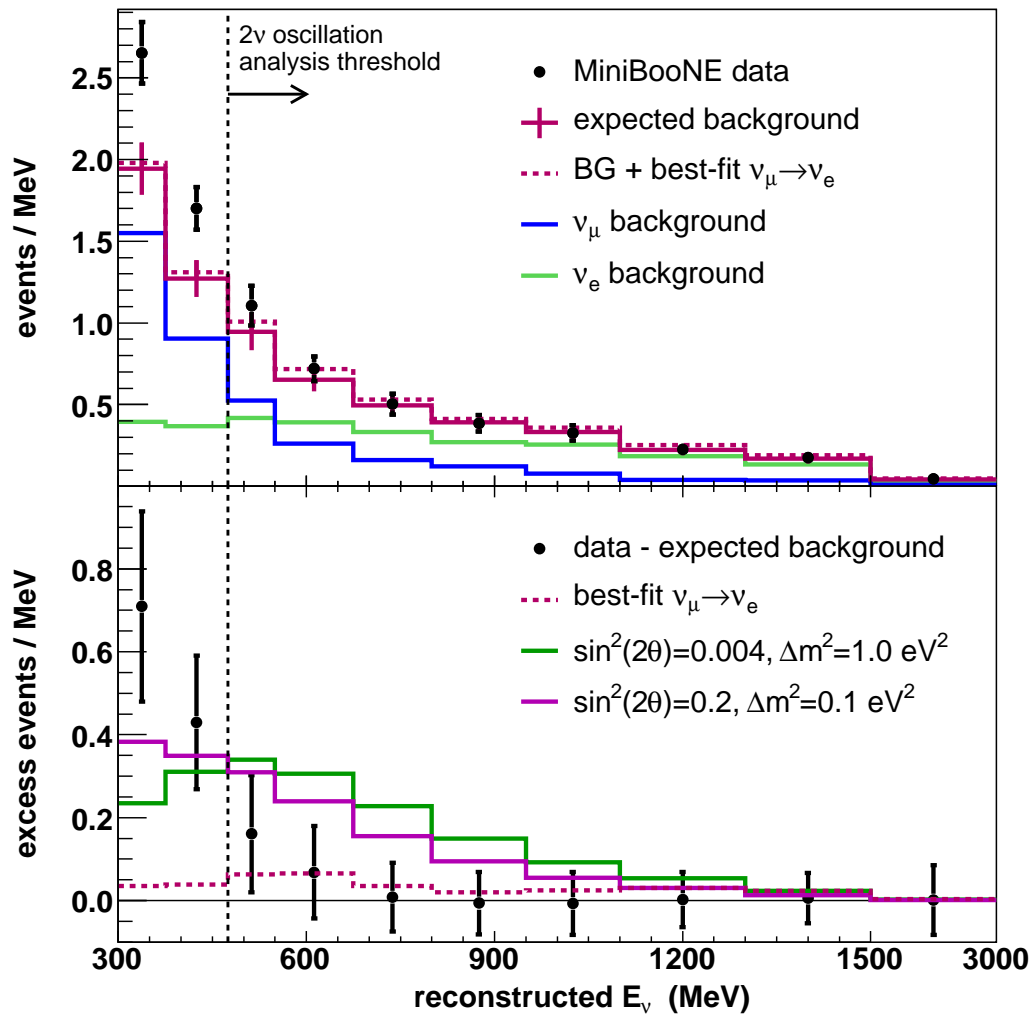
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- Experimental results are frequently wrong!

2c: MiniBooNE Anomaly

Search for ν_e appearance in ν_μ beam (from π^+ decay in flight) (ref: arXiv:0704.1500)

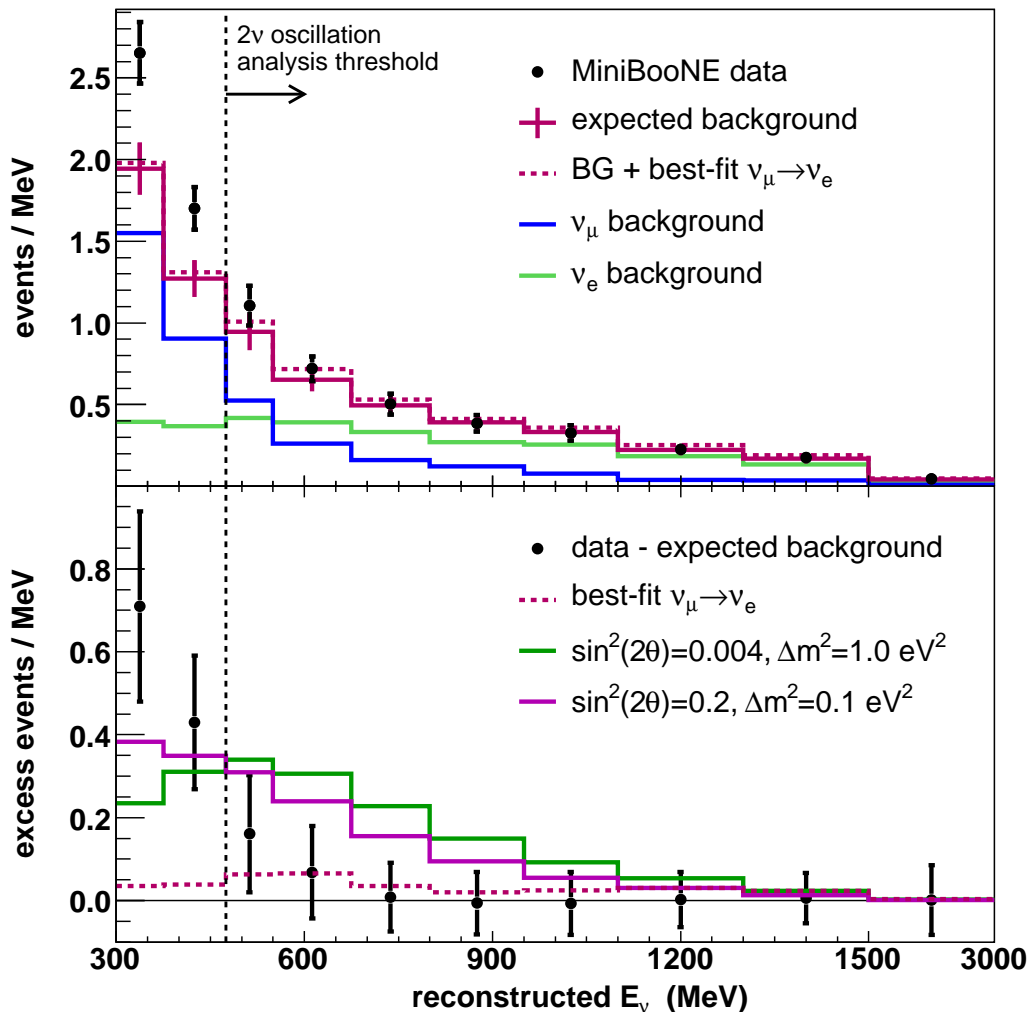
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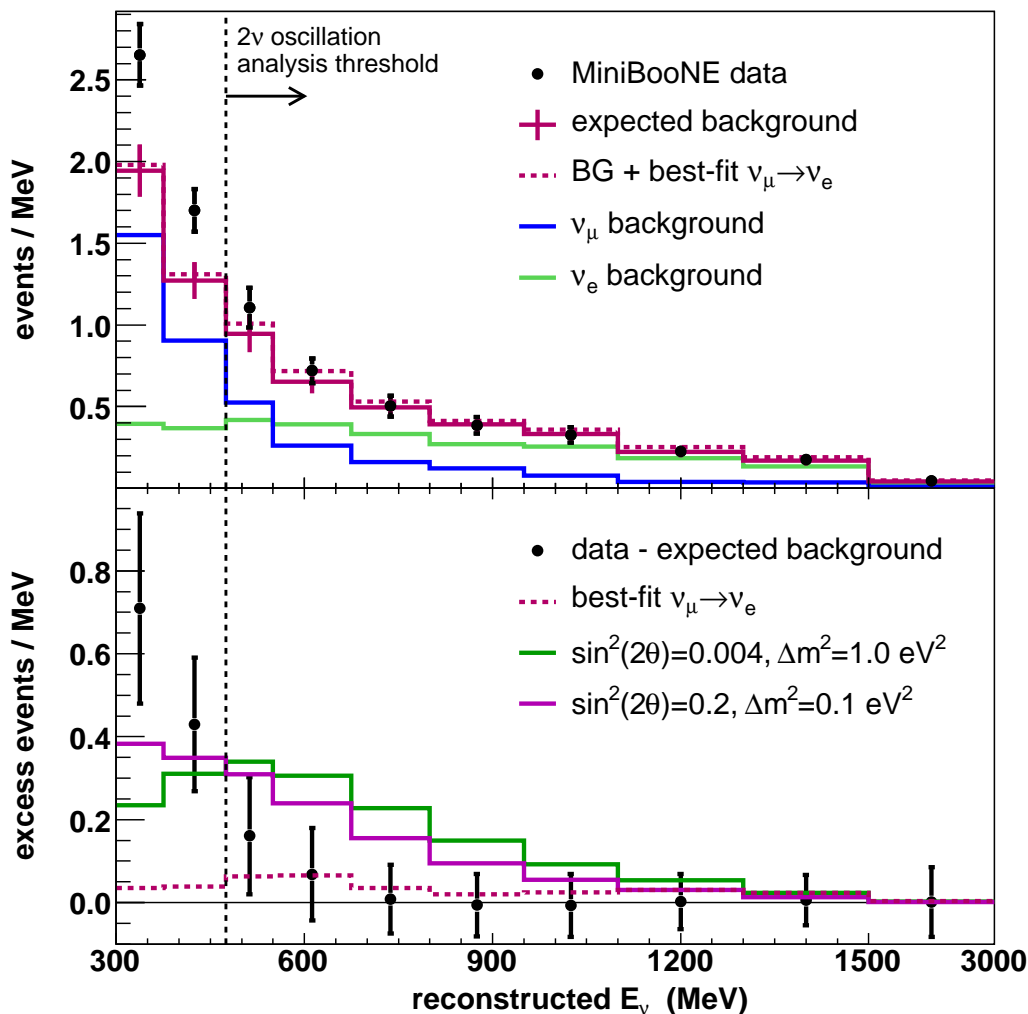
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$96 \pm 17 \pm 20$ excess events below 475 MeV!

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- Excess probably due to low-energy hadronic/nuclear effects (e.g. R.J. Hill, arXiv:0905.0291)

2d: CDF Ghosts arXiv:0810.5357

Look at di-muon sample:

$p_T(\mu) > 3 \text{ GeV}$, $|\eta(\mu)| < 0.7$, $5 \text{ GeV} < M_{\mu\mu} < 80 \text{ GeV}$,
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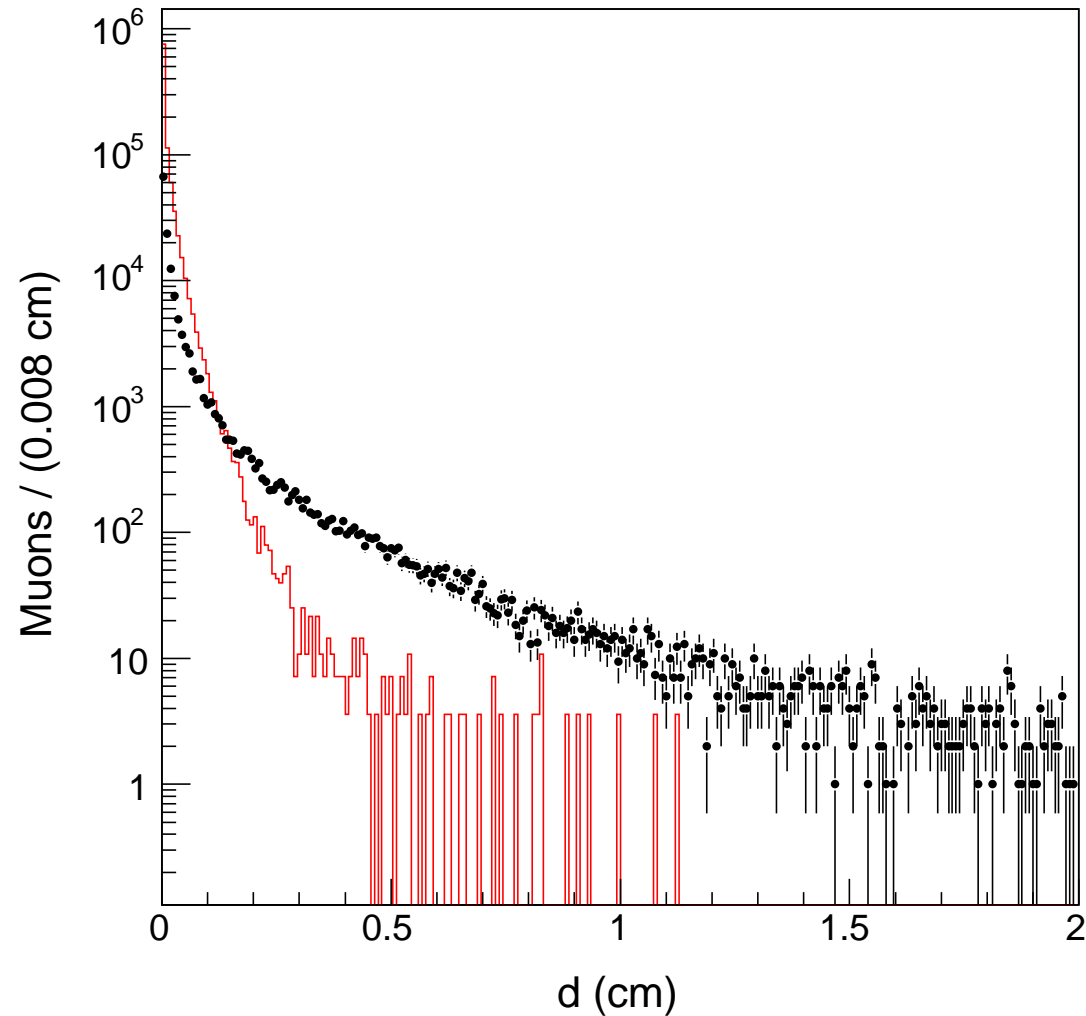
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W/ “loose SVX”, requiring origin w/in 10 cm of primary
 vertex: half of “ghosts” gone; $\epsilon_{\text{SM}} = 88\%$.

CDF Ghosts (cont'd)

Many ghost events have very large impact parameter:



CDF Ghost Events with Additional Muons

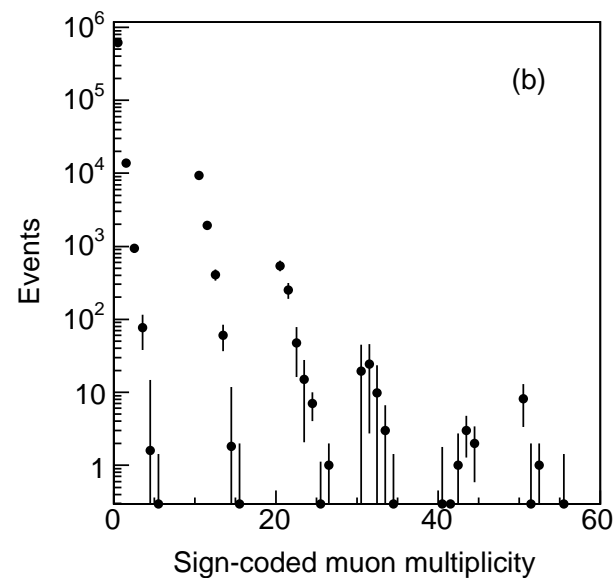
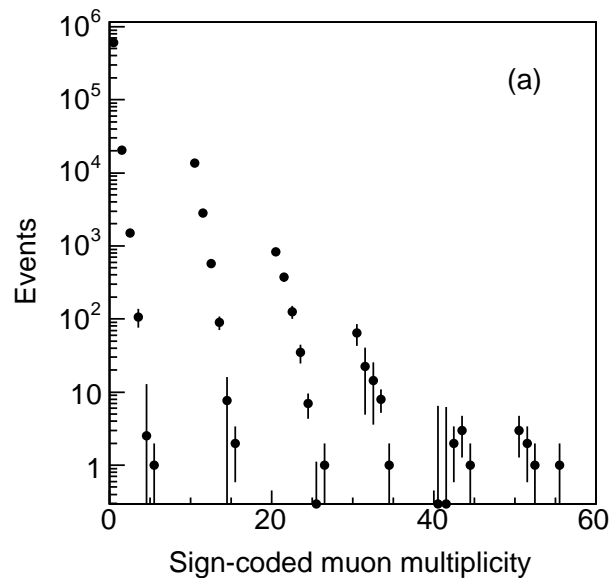
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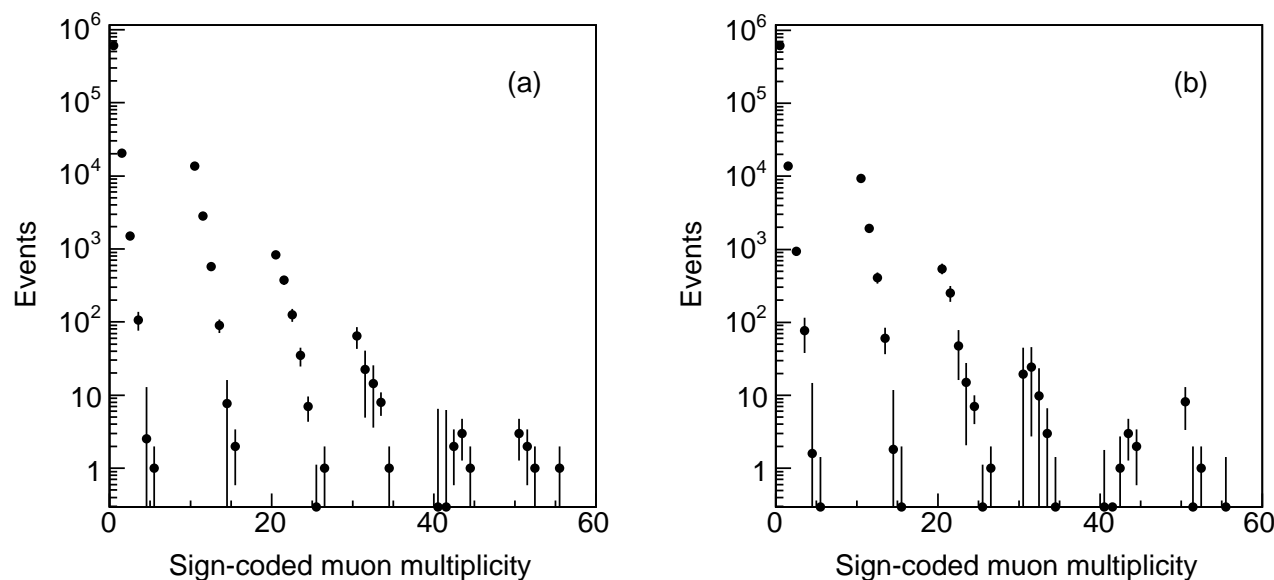
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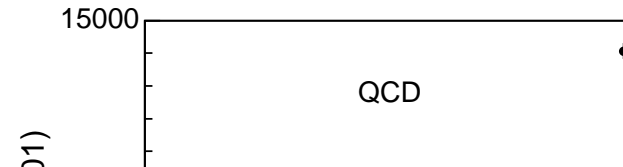
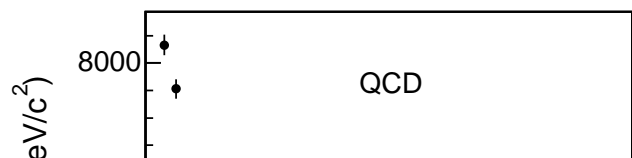
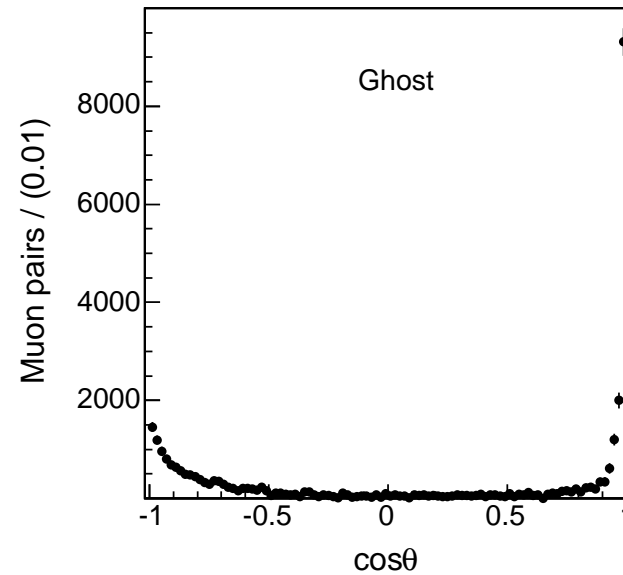
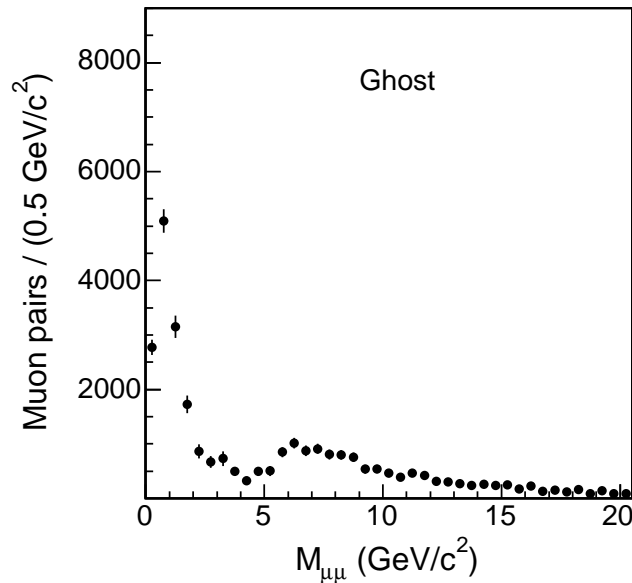
Nearly equal numbers of same-sign and opposite-sign combinations!

CDF Ghost Events: Kinematics

- *No info on kinematic distribution of primary μ 's in ghost sample!!?*

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- No info on kinematic distribution of primary μ 's in ghost sample!!?
- $M_{\mu_p\mu_s}$ distribution, and large event no., suggest low mass scale of ghost events!



CDF Ghosts: Checks

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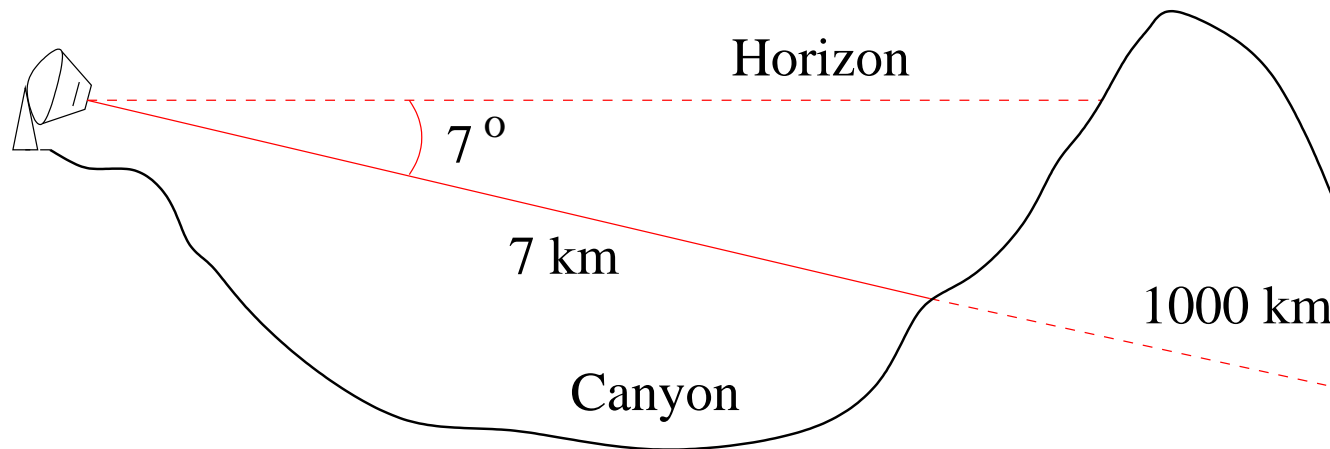
- D0 data (obviously)
- Low- E data: Look for events with muon(s) at large impact parameter or from far distant decay vertex. Scales are centimeter: Don't need fancy SVX detector!
pp at RHIC; CERN SpS data? FNAL fixed target; HERA-B; ...

2e: SHALON Anomaly arXiv:0903.4654

SHALON: 11.2 m² Cherenkov telescope in Tien–Shan mountains (China), 3.3 km a.s.l: Can see “below the horizon:”

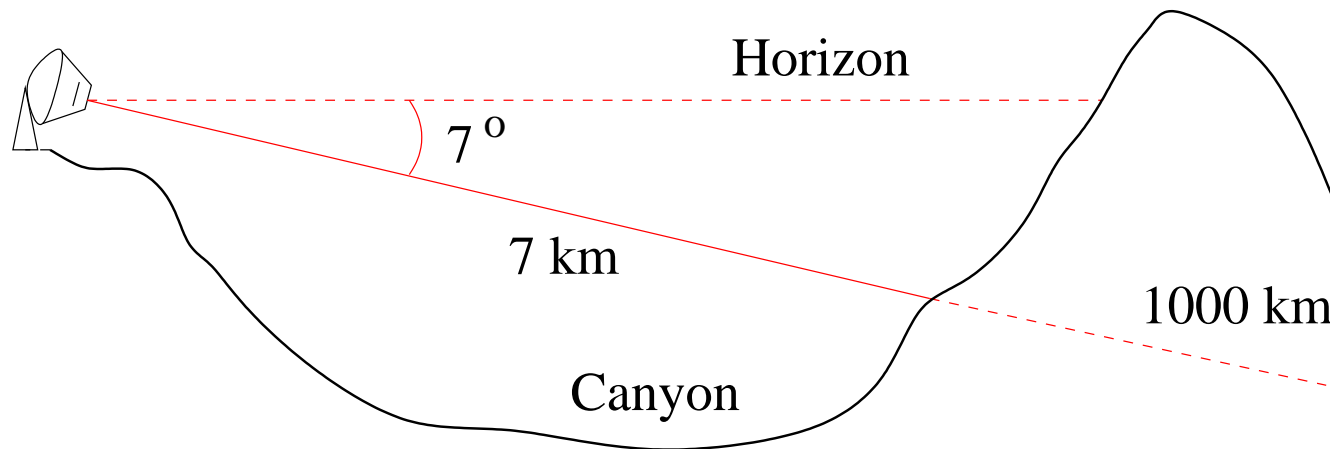
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234 h observation time at 7° below horizon

SHALON Anomaly (cont'd)

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- Their interpretation: **500 MeV neutrino** decaying radiatively, $\nu_h \rightarrow \nu \gamma$, with $c\tau_{\nu_h} \lesssim 100 \text{ m}$ (i.e. $\gamma c\tau_{\nu_h} \lesssim 1000 \text{ km}$)

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- Same model supposedly explains MiniBooNE anomaly (arXiv:0902.3802)

SHALON Anomaly: comments

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- Something for future expts of this kind to get excited, or worry, about.

3 Challenges to the Concordance Model

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Is this evidence for MOND/TeVeS?

Cannot make this claim! Nobody knows how structure formation works in MOND!

CDM Problems: To-do-list

- **Check/improve simulations:** effect of baryons, black holes, ...

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- **Improve measurements of galaxy kinematics:**
 - Better measurements of rotation curves of other galaxies
 - **GAIA** for our own galaxy. Launch 2011. Should measure 10^9 stars in 3D (position and radial velocity); **get tangential velocity in 40M stars to better than 0.5 km/s!** $v_{\text{rot}} \sim 220$ km/s.

3b: CMB Anomalies

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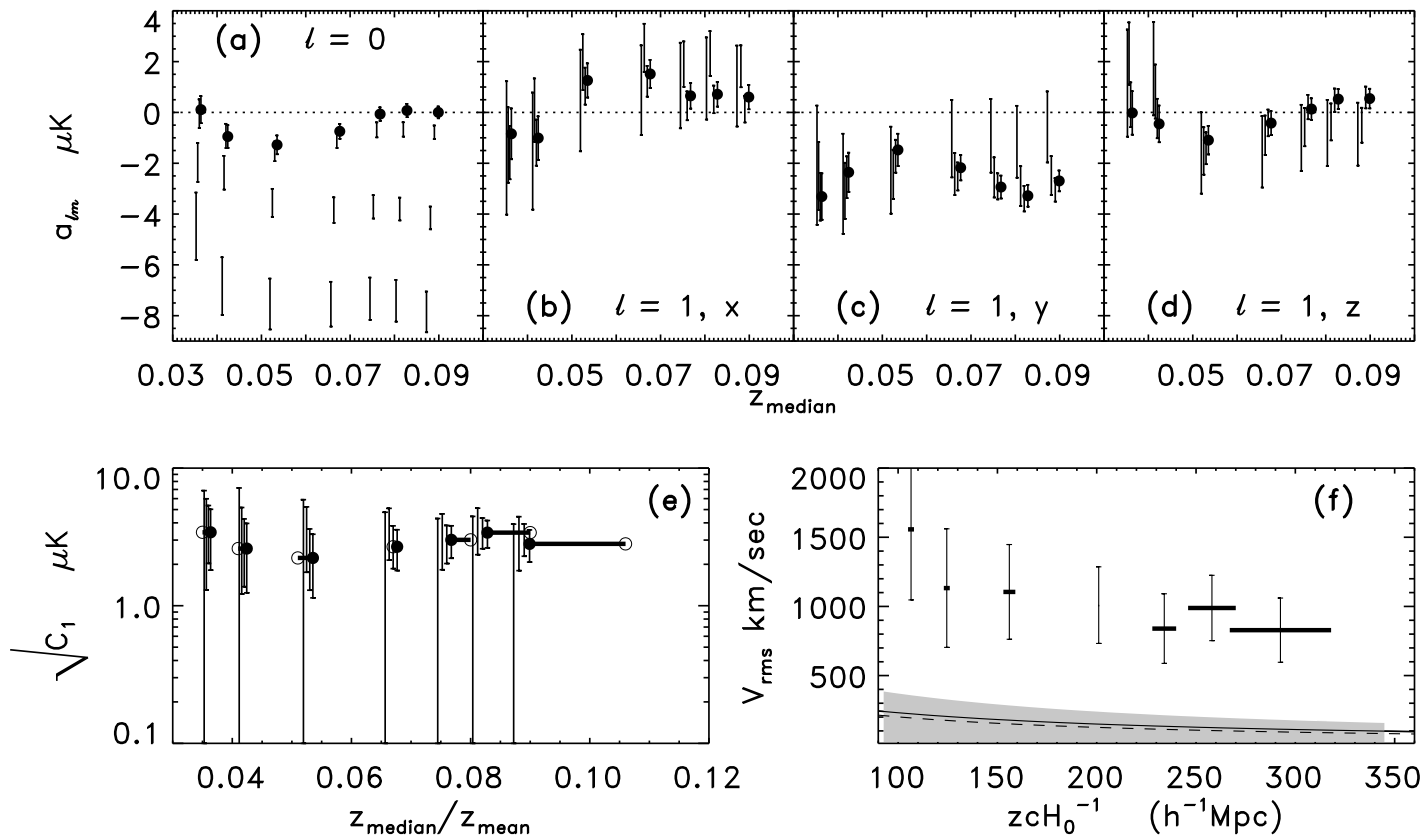
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- But: Any measured complicated distribution is bound to look unlikely in some ways! (Any one phase space point has vanishing probability.)
- WMAP still dominates determinations of many cosmological parameters.
- PLANCK launched successfully May 14. Should improve precision of parameters, e.g. $\Omega_{\text{DM}}h^2$ to few%: Also need improved calculations!

3c: Large Scale Velocity Flows

Kashlinsky et al., [arXiv:0809.3734](https://arxiv.org/abs/0809.3734): Measure peculiar velocities of clusters of galaxies through Sunyaev-Zeldovich effect: CMB photons scatter off hot plasma.

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- **Sign for DGP gravity?** Graviton gets tiny mass, i.e. longitudinal component; *increases* density perturbations at small scales. **Afshordi et al., arXiv:0812.4684**

3d: Λ Problem

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- Other way round: testing details of inflation (e.g. non-Gaussianity; gravity waves) may shed new light on Λ problem! PLANCK, again!

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- Large scale velocity flows mostly ignored so far: ~ 20 citations; also true for most other anomalies I discussed.